• Bending magnet review

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Homework Assignment #01: Chapter Chapter 2: 2,3,5,6,8 due Thursday, January 29, 2015



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The Fourier transform of this pulse is the spectrum of the radiation from the bending magnet.

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$$\mathcal{E}_c[\text{keV}] = 0.665 \mathcal{E}^2[\text{GeV}]B[\text{T}]$$

When the radiation pulse time is Fourier transformed, we obtain the spectrum of a bending magnet.



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Scaling by the characteristic energy, gives a universal curve



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$$1.33 \times 10^{13} \mathcal{E}^2 I\left(\frac{\omega}{\omega_c}\right)^2 K_{2/3}^2\left(\frac{\omega}{2\omega_c}\right)$$

where $K_{2/3}$ is a modified Bessel function of the second kind.



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We can calculate this for the ESRF where $\mathcal{E} = 6$ GeV, B = 0.8 T, $\mathcal{E}_c = 19.2$ keV and the bending radius $\rho = 24.8$ m. Assuming that the aperture is 1 mm² at a distance of 20 m, the angular aperture is 1/20 = 0.05 mrad and the flux at the characteristic energy is given by:

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 $\mathsf{Flux} = (1.95 \times 10^{13})(0.05^2 \mathsf{mrad}^2)(6^2 \mathsf{GeV}^2)(0.2 \mathsf{A}) = 3.5 \times 10^{11} \mathsf{ph/s}/0.1\% \mathsf{BW}$

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A bending magnet also produces circularly polarized radiation



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The result is circularly polarized radiation above and below the on-axis radiation.

Wiggler



Wiggler



Like bending magnet except:

Wiggler



Like bending magnet except:

• larger $\vec{B} \rightarrow$ higher E_c

Wiggler



Like bending magnet except:

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- more bends \rightarrow higher power



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Different from bending magnet:

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Different from bending magnet:

- shallow bends \rightarrow smaller source



Like bending magnet except:

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Different from bending magnet:

- shallow bends \rightarrow smaller source
- interference \rightarrow peaked spectrum

• The electron's trajectory through a wiggler can be considered as a series of short circular arcs, each like a bending magnet

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- This results in a significantly higher power load on all downstream components

```
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```



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From the electron trajectory:

$$x = A\sin\left(k_u z\right)$$



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Define a dimensionless quantity, K which scales $\alpha_{\rm max}$ to the natural opening angle of the radiation, $1/\gamma$

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$$K = \alpha_{max} \gamma$$

C. Segre (IIT)

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$$\rho^2 = [x + (\rho - A)]^2 + z^2$$



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$$= \lambda_{u} \left(1 + \frac{A^{2}k_{u}^{2}}{4} \right) = \lambda_{u} \left(1 + \frac{1}{4}\frac{K^{2}}{\gamma^{2}} \right)$$
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The emitted wave travels slightly faster than the electron.

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Consider an electron traveling through the undulator and emitting radiation at the first maximum excursion from the center.



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$$\lambda_1 = cT' - \lambda_u$$

PHYS 570 - Spring 2015

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The emitted wave travels slightly faster than the electron.

It moves cT' in the time the electron travels a distance λ_u along the undulator.

The observer sees radiation with a compressed wavelength, along with harmonics which satisfy the same condition.

$$n\lambda_n = cT' - \lambda_u$$

PHYS 570 - Spring 2015

The fundamental wavelength must be corrected for the observer angle θ

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The fundamental wavelength must be corrected for the observer angle θ

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$$=\lambda_u \left(\mathbf{S} \frac{\mathbf{c}}{\mathbf{v}} - \cos \theta \right)$$

$$=\lambda_u\left(\left[1+\frac{\kappa^2}{4\gamma^2}\right]\frac{1}{\beta}-\cos\theta\right)$$

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$${m S}pprox 1+{{m K}^2\over 4\gamma^2}$$

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Over the time T' the electron actually travels a distance $S\lambda_u$, so that

$$T' = \frac{S\lambda_u}{v}$$

 $\frac{K^2}{4\gamma^2}$

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2

$$=\lambda_{\mu}\left(\left[1+\frac{\kappa^{2}}{4\gamma^{2}}\right]\frac{1}{\beta}-\cos\theta\right)\qquad \qquad S\approx1+\frac{\kappa^{2}}{4\gamma^{2}}$$

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$$\lambda_1 \approx \lambda_u \left(\frac{1}{\beta} + \frac{\kappa^2}{4\gamma^2\beta} - 1 + \frac{\theta^2}{2} \right) = \frac{\lambda_u}{2\gamma^2} \left(\frac{2\gamma^2}{\beta} + \frac{\kappa^2}{2\beta} - 2\gamma^2 + \gamma^2\theta^2 \right)$$

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regrouping terms
$$\lambda_{1} \approx \frac{\lambda_{u}}{2\gamma^{2}} \left(\frac{2\gamma^{2}}{\beta} + \frac{\kappa^{2}}{2\beta} - 2\gamma^{2} + \gamma^{2}\theta^{2} \right)$$
$$\approx \frac{\lambda_{u}}{2\gamma^{2}} \left(2\gamma^{2} \left[\frac{1}{\beta} - 1 \right] + \frac{\kappa^{2}}{2\beta} - (\gamma\theta)^{2} \right)$$

regrouping terms

$$\lambda_1 pprox rac{\lambda_u}{2\gamma^2} \left(rac{2\gamma^2}{eta} + rac{\kappa^2}{2eta} - 2\gamma^2 + \gamma^2 heta^2
ight)$$
 $pprox rac{\lambda_u}{2\gamma^2} \left(2\gamma^2 \left[rac{1}{eta} - 1
ight] + rac{\kappa^2}{2eta} - (\gamma heta)^2
ight)$

$$\gamma = \sqrt{\frac{1}{1 - \beta^2}}$$

$$\begin{split} \lambda_{1} &\approx \frac{\lambda_{u}}{2\gamma^{2}} \left(\frac{2\gamma^{2}}{\beta} + \frac{\kappa^{2}}{2\beta} - 2\gamma^{2} + \gamma^{2}\theta^{2} \right) \\ &\approx \frac{\lambda_{u}}{2\gamma^{2}} \left(2\gamma^{2} \left[\frac{1}{\beta} - 1 \right] + \frac{\kappa^{2}}{2\beta} - (\gamma\theta)^{2} \right) \\ &\approx \frac{\lambda_{u}}{2\gamma^{2}} \left(2\gamma^{2} \left[\frac{1}{\beta} - 1 \right] + \frac{\kappa^{2}}{2\beta} - (\gamma\theta)^{2} \right) \\ &\approx \frac{\lambda_{u}}{2\gamma^{2}} \left(2\frac{1}{1 - \beta^{2}} \left[\frac{1 - \beta}{\beta} \right] + \frac{\kappa^{2}}{2\beta} - (\gamma\theta)^{2} \right) \end{split}$$

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$$\begin{split} _{1} &\approx \frac{\lambda_{u}}{2\gamma^{2}} \left(\frac{2\gamma^{2}}{\beta} + \frac{K^{2}}{2\beta} - 2\gamma^{2} + \gamma^{2}\theta^{2} \right) \\ &\approx \frac{\lambda_{u}}{2\gamma^{2}} \left(2\gamma^{2} \left[\frac{1}{\beta} - 1 \right] + \frac{K^{2}}{2\beta} - (\gamma\theta)^{2} \right) \\ &\approx \frac{\lambda_{u}}{2\gamma^{2}} \left(2\gamma^{2} \left[\frac{1}{\beta} - 1 \right] + \frac{K^{2}}{2\beta} - (\gamma\theta)^{2} \right) \\ &\approx \frac{\lambda_{u}}{2\gamma^{2}} \left(2\frac{1}{1 - \beta^{2}} \left[\frac{1 - \beta}{\beta} \right] + \frac{K^{2}}{2\beta} - (\gamma\theta)^{2} \right) \\ &\approx \frac{\lambda_{u}}{2\gamma^{2}} \left(\frac{2}{\beta(1 + \beta)} + \frac{K^{2}}{2\beta} - [\gamma\theta]^{2} \right) \\ \end{split}$$

λ

$$\lambda_1 \approx \frac{\lambda_u}{2\gamma^2} \left(\frac{2}{\beta(1+\beta)} + \frac{\kappa^2}{2\beta} - (\gamma\theta)^2 \right)$$

If we assume that $\beta \sim 1$ for these highly relativistic electrons

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for a typical undulator $\gamma \sim 10^4$, $K \sim 1$, and $\lambda_u \sim 2$ cm so we estimate

$$\lambda_1 \approx \frac{2 \times 10^{-2}}{2 \ (10^4)^2} \left(1 + \frac{(1)^2}{2}\right)$$

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Recall that we developed an expression for the Doppler time compression of the emission from a moving electron as a function of the observer angle.



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Recall that we developed an expression for the Doppler time compression of the emission from a moving electron as a function of the observer angle.

This can be rewritten in terms of the coordinates in the figure using the vector of unit length in the observer direction:

$$\vec{\mathbf{n}} = \left\{ \phi, \psi, \sqrt{1 - \theta^2} \right\}$$
$$\vec{\beta} = \beta \left\{ \alpha, \mathbf{0}, \sqrt{1 - \alpha^2} \right\}$$



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$$ec{n} pprox \left\{ \phi, \psi, (1 - heta^2/2)
ight\}$$

 $ec{eta} pprox eta \left\{ lpha, 0, (1 - lpha^2/2)
ight\}$



$$egin{split} rac{dt}{dt'} &= 1 - ec{n} \cdot ec{eta}(t') \ &pprox 1 - eta \left[lpha \phi + \left(1 - rac{ heta^2}{2} - rac{lpha^2}{2}
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$$\begin{split} \frac{dt}{dt'} &= 1 - \vec{n} \cdot \vec{\beta}(t') & \vec{n} \approx \left\{\phi, \psi, (1 - \theta^2/2)\right\} \\ &\approx 1 - \beta \left[\alpha \phi + \left(1 - \frac{\theta^2}{2} - \frac{\alpha^2}{2}\right)\right] & \vec{\beta} \approx \beta \left\{\alpha, 0, (1 - \alpha^2/2)\right\} \\ &\frac{dt}{dt'} \approx 1 - \left(1 - \frac{1}{2\gamma^2}\right) \left(1 + \alpha \phi - \frac{\theta^2}{2} - \frac{\alpha^2}{2}\right) \end{split}$$

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$$rac{dt}{dt'}pprox 1 - \left(1 - rac{1}{2\gamma^2}
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This differential equation can be solved, realizing that ϕ and θ are constant while $\alpha(t')$ varies as the electron moves through the insertion device, and gives:

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 $\omega_1 t = \omega_u t'$

 $\omega_1\gg\omega_u$ as expected because of the Doppler compression

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$$\omega_1 t = \omega_u t' - \frac{\kappa^2/4}{1 + (\gamma \theta)^2 + \kappa^2/2} \sin(2\omega_u t')$$

 $\omega_1\gg\omega_u$ as expected because of the Doppler compression , but they are not proportional because of the second

$$\begin{split} \frac{dt}{dt'} &\approx 1 - \left(1 - \frac{1}{2\gamma^2}\right) \left(1 + \alpha\phi - \frac{\theta^2}{2} - \frac{\alpha^2}{2}\right) \\ &\approx 1 - 1 - \alpha\phi + \frac{\theta^2}{2} + \frac{\alpha^2}{2} + \frac{1}{2\gamma^2} = \frac{1}{2} \left(\theta^2 + \alpha^2 + \frac{1}{\gamma^2}\right) - \alpha\phi \end{split}$$

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 $\omega_1\gg\omega_u$ as expected because of the Doppler compression , but they are not proportional because of the second and third terms.

$$\begin{split} \frac{dt}{dt'} &\approx 1 - \left(1 - \frac{1}{2\gamma^2}\right) \left(1 + \alpha\phi - \frac{\theta^2}{2} - \frac{\alpha^2}{2}\right) \\ &\approx 1 - 1 - \alpha\phi + \frac{\theta^2}{2} + \frac{\alpha^2}{2} + \frac{1}{2\gamma^2} = \frac{1}{2} \left(\theta^2 + \alpha^2 + \frac{1}{\gamma^2}\right) - \alpha\phi \end{split}$$

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The motion of the electron, $\sin \omega_u t'$, is always sinusoidal, but because of the additional terms, the motion as seen by the observer, $\sin \omega_1 t$, is not.

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$$\omega_1 t = \omega_u t' - \frac{\kappa^2/4}{1 + (\gamma \theta)^2 + \kappa^2/2} \sin(2\omega_u t')$$

Suppose we have K = 1 and $\theta = 0$ (on axis), then

$$\omega_1 t = \omega_u t' - \frac{\kappa^2/4}{1 + (\gamma \theta)^2 + \kappa^2/2} \sin(2\omega_u t')$$

Suppose we have K = 1 and $\theta = 0$ (on axis), then

$$\omega_1 t = \omega_u t' + \frac{1}{6} \sin\left(2\omega_u t'\right)$$

$$\omega_{1}t = \omega_{u}t' - \frac{K^{2}/4}{1 + (\gamma\theta)^{2} + K^{2}/2} \sin(2\omega_{u}t')$$
Suppose we have $K = 1$ and $\theta = 0$
(on axis), then
 $\omega_{1}t = \omega_{u}t' + \frac{1}{6}\sin(2\omega_{u}t')$
Plotting $sin\omega_{u}t'$ and $sin\omega_{1}t$ shows the deviation from sinusoidal.

$$I = \bigcup_{0} \bigcup_{0} \bigcup_{\pi/2} \bigcup_{\pi/2}$$

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Similarly, for $K = 2$

Phase Angle (radians)

$$\omega_1 t = \omega_u t' - \frac{K^2/4}{1 + (\gamma \theta)^2 + K^2/2} \sin(2\omega_u t')$$

Suppose we have $K = 1$ and $\theta = 0$
(on axis), then

$$\omega_1 t = \omega_u t' + \frac{1}{6} \sin\left(2\omega_u t'\right)$$

Plotting $sin\omega_u t'$ and $sin\omega_1 t$ shows the deviation from sinusoidal.

Similarly, for K = 2 and K = 5, the deviation becomes more pronounced.



$$\omega_1 t = \omega_u t' - \frac{\kappa^2/4}{1 + (\gamma \theta)^2 + \kappa^2/2} \sin\left(2\omega_u t'\right)$$

Suppose we have K = 1 and $\theta = 0$ (on axis), then

$$\omega_1 t = \omega_u t' + \frac{1}{6} \sin\left(2\omega_u t'\right)$$

Plotting $sin\omega_u t'$ and $sin\omega_1 t$ shows the deviation from sinusoidal.

Similarly, for K = 2 and K = 5, the deviation becomes more pronounced. This shows how higher harmonics must be present in the radiation as seen by the observer.




$$\omega_{1}t = \omega_{u}t' - \frac{K^{2}/4}{1 + (\gamma\theta)^{2} + K^{2}/2} \sin(2\omega_{u}t') - \frac{2K\gamma}{1 + (\gamma\theta)^{2} + K^{2}/2} \phi \sin(\omega_{u}t')$$

$$\int_{0.5}^{1} \frac{K=2}{\theta=0} \quad \text{When } K = 2 \text{ and } \theta = \phi = 1/\gamma, \text{ we have}$$

$$\omega_{1}t = \omega_{u}t' + \frac{1}{4}\sin(2\omega_{u}t') + \sin\omega_{u}t'$$

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When K = 2 and $\theta = \phi = 1/\gamma$, we have

$$\omega_1 t = \omega_u t' + \frac{1}{4} \sin(2\omega_u t') + \sin\omega_u t'$$

The last term introduces an antisymmetric term which skews the function

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When ${\cal K}=2$ and $\theta=\phi=1/\gamma$, we have

$$\omega_1 t = \omega_u t' + \frac{1}{4} \sin(2\omega_u t') + \sin\omega_u t'$$

The last term introduces an antisymmetric term which skews the function and leads to the presence of forbidden harmonics $(2^{nd}, 4^{th}, \text{ etc})$ in the radiation from the undulator compared to the on-axis radiation.



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• Brilliance is 6 orders larger than a bending magnet



- Brilliance is 6 orders larger than a bending magnet
- Both odd and even harmonics appear



- Brilliance is 6 orders larger than a bending magnet
- Both odd and even harmonics appear
- Harmonics can be tuned in energy (dashed lines)