## Today's Outline - January 20, 2015

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- Coherence of x-ray sources


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- Coherence of $x$-ray sources
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Homework Assignment \#01:
Chapter Chapter 2: 2,3,5,6,8
due Thursday, January 29, 2015

## Coherence: what is it?

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- they do not travel in a perfectly co-linear direction

Because of these imperfections the "coherence length" of an $x$-ray beam is finite and we can calculate it.

## Longitudinal coherence

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## Transverse coherence

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$$
L_{T}=\frac{\lambda R}{2 D}
$$

## Coherence lengths at the APS

For a typical $3^{\text {rd }}$ generation undulator source, such as at the Advanced Photon Source the vertical source size is $D=100 \mu \mathrm{~m}$ and we are typically $R=50 \mathrm{~m}$ away with our experiment. If we assume a typical wavelength of $\lambda=1 \AA$, and a monochromator resolution of $\Delta \lambda / \lambda=10^{-5}$ we have for the vertical direction:

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L_{T}=\frac{\lambda R}{2 D}=\frac{\left(1 \times 10^{-10}\right) \cdot 50}{2 \cdot\left(100 \times 10^{-6}\right)}=25 \mu \mathrm{~m}
\end{gathered}
$$

## X-ray tube schematics

Fixed anode tube


- low power
- low maintenance


## X-ray tube schematics

Fixed anode tube


- low power
- low maintenance

Rotating anode tube


- high power
- high maintenance


## X-ray tube spectrum

Mini-X Output X-Ray Spectrum: Ag Target @ 40 kV


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- Minimum wavelength (maximum energy) set by accelerating potential


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- Highest intensity at the characteristic fluorescence emission energy of the anode material


## X-ray tube spectrum

Mini-X Output X-Ray Spectrum: Ag Target @ 40 kV


- Minimum wavelength (maximum energy) set by accelerating potential
- Bremßtrahlung radiation provides smooth background (charged particle deceleration)
- Highest intensity at the characteristic fluorescence emission energy of the anode material
- Unpolarized, incoherent x-rays emitted in all directions from anode surface, must be collimated with slits


## Synchrotron sources

## Bending magnet



## Synchrotron sources

## Bending magnet



- Wide horizontal beam


## Synchrotron sources

## Bending magnet



- Wide horizontal beam
- Broad spectrum to high energies


## Synchrotron sources

Bending magnet


Undulator


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## Synchrotron sources

Bending magnet


- Wide horizontal beam
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Undulator


- Highly collimated beam


## Synchrotron sources

Bending magnet


- Wide horizontal beam
- Broad spectrum to high energies

Undulator


- Highly collimated beam
- Highly peaked spectrum with harmonics


## Bending magnet spectra



## Bending magnet spectra



Lower energy sources, such as NSLS have lower peak energy and higher intensity at the peak.

## Bending magnet spectra



Lower energy sources, such as NSLS have lower peak energy and higher intensity at the peak.
Higher energy sources, such as APS have higher energy spectrum and are only off by a factor of 2 intensity at low energy.

## Bending magnet spectra



## Bending magnet spectra



Logarithmic scale shows clearly how much more energetic and intense the bending magnet sources at the 6 GeV and 7 GeV sources are.

## Review of special relativity

## $\vec{V}$

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\beta=\frac{v}{c}
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\beta=\frac{v}{c} \quad \gamma=\sqrt{\frac{1}{1-\beta^{2}}} \\
E=\gamma m c^{2} \\
\beta=\sqrt{1-\frac{1}{\gamma^{2}}} \longrightarrow \beta \approx 1-\frac{1}{2} \frac{1}{\gamma^{2}}
\end{gathered}
$$

use binomial expansion since $1 / \gamma^{2} \ll 1$

## Review of special relativity

Let's calculate these quantities for an electron at NSLS and APS

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& \\
& \gamma=\frac{1.5 \times 10^{9}}{0.511 \times 10^{6}}=2935
\end{array}
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## "Headlight" effect

In electron rest frame:

emission is symmetric about the axis of the acceleration vector

## "Headlight" effect

In electron rest frame:

emission is symmetric about the axis of the acceleration vector

In lab frame:

emission is pushed into the direction of motion of the electron

## Relativistic emission


the electron is in constant transverse acceleration due to the Lorentz force from the magnetic field of the bending magnet

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\vec{F}=e \vec{v} \times \vec{B}=m_{e} \vec{a}
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E_{\max } \approx \gamma^{3} \omega_{0}
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for the APS, with $\gamma \approx 10^{4}$ we have

$$
E_{\max } \approx\left(10^{4}\right)^{3} \cdot 10^{6}=10^{18}
$$

## Flux and brilliance

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All these quantities are conveniently taken into account in a measure called brilliance

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## Computing brilliance

$$
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\alpha \approx x / z \quad \beta \approx y / z
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where $z$ is the distance from the source over which there is a lateral spread $x$ and $y$ in each direction

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called the time compression ratio.

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The Doppler shift is defined in terms of the time compression ratio

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The observer, looking in the plane of the circular trajectory, "sees" the electron oscillate over a half period in a time $\Delta t$ (observer's frame). The electron, in the laboratory frame, travels this arc in:

$$
\Delta t^{\prime}=\frac{(1 / \gamma) \rho}{v}=\frac{1}{\gamma \omega_{0}}
$$

Because of the Doppler shift, the observer sees the electron emitting a pulse of radiation of length

$$
\Delta t \propto \frac{\Delta t^{\prime}}{\gamma^{2}}=\frac{1}{\gamma^{3} \omega_{o}}
$$

The Fourier transform of this pulse is the spectrum of the radiation from the bending magnet.

## Characteristic Energy of a Bending Magnet

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$$
\mathcal{E}_{c}[\mathrm{keV}]=0.665 \mathcal{E}^{2}[\mathrm{GeV}] B[\mathrm{~T}]
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## Bending magnet spectrum

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Scaling by the characteristic energy, gives a universal curve
$1.33 \times 10^{13} \mathcal{E}^{2} I\left(\frac{\omega}{\omega_{c}}\right)^{2} K_{2 / 3}^{2}\left(\frac{\omega}{2 \omega_{c}}\right)$
where $K_{2 / 3}$ is a modified Bessel function of the second kind.


